

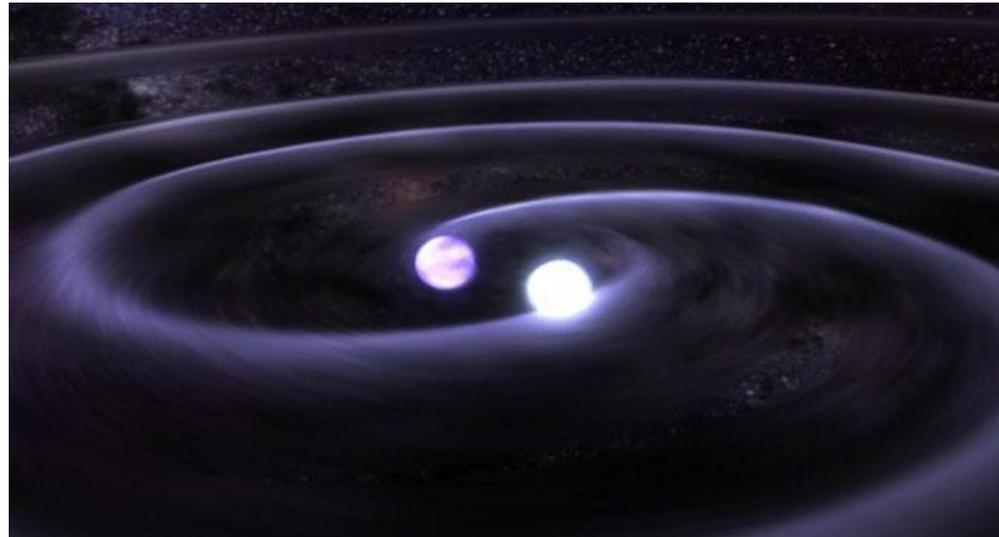
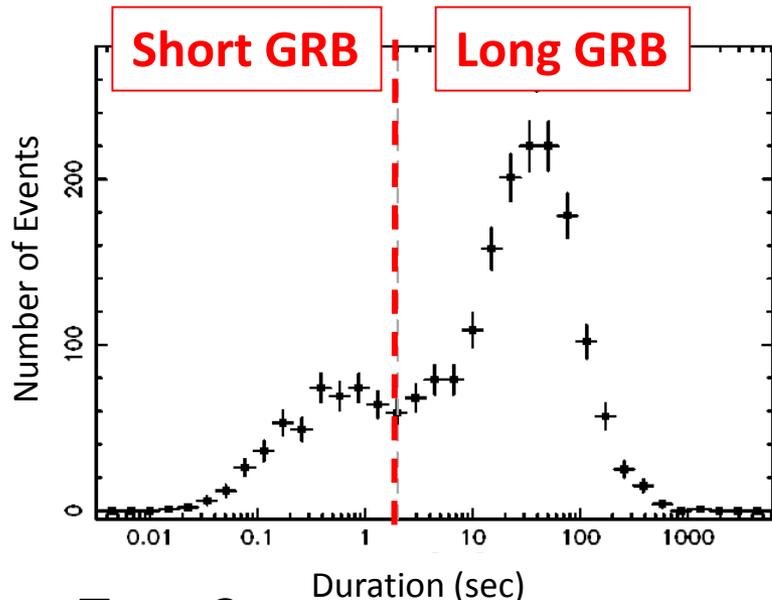
# Development of Silicon-based X-ray Transient Monitor onto a Microsat

Kanazawa University

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# Short gamma-ray bursts



credit NASA

- $T_{90} < 2 \text{ sec}$
- Associated with the coalescence of a binary system of compact objects(?)

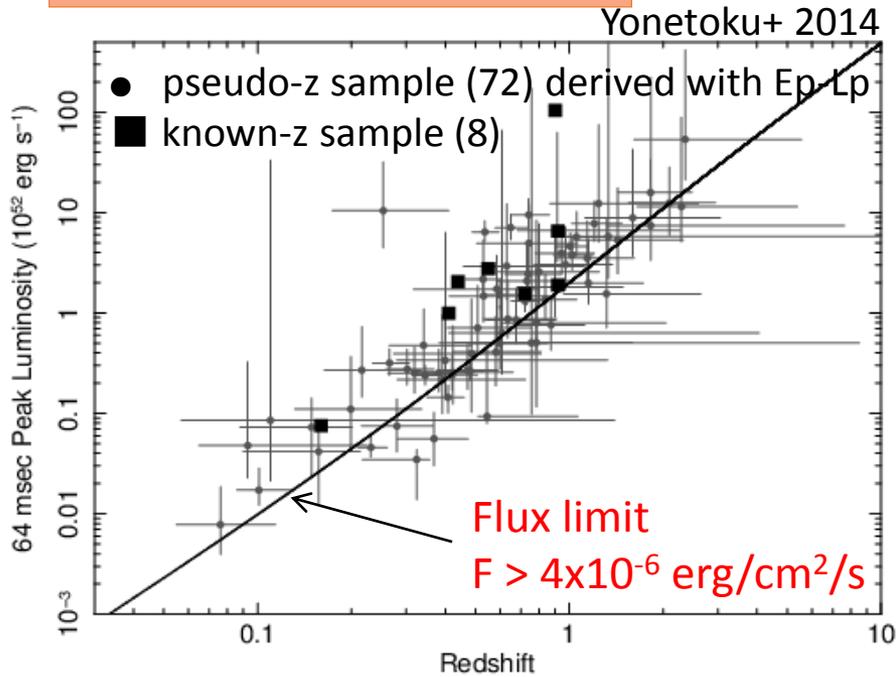
*What would be the keys to understanding of the progenitor?*

- gravitational-wave transients coincident with Short GRBs
  - GW? → KAGRA, advanced LIGO, & advanced VIRGO (2018-)
  - GRB? Swift/BAT(Hard X-ray band), Fermi/GBM (wide band, wide FoV, but moderate localization), and ...

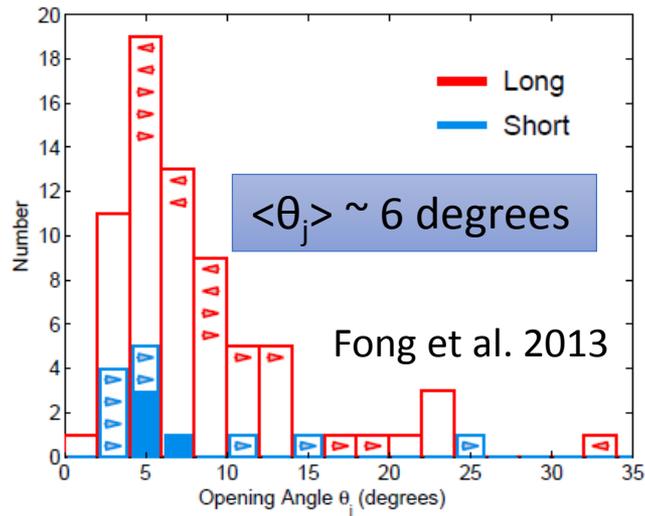
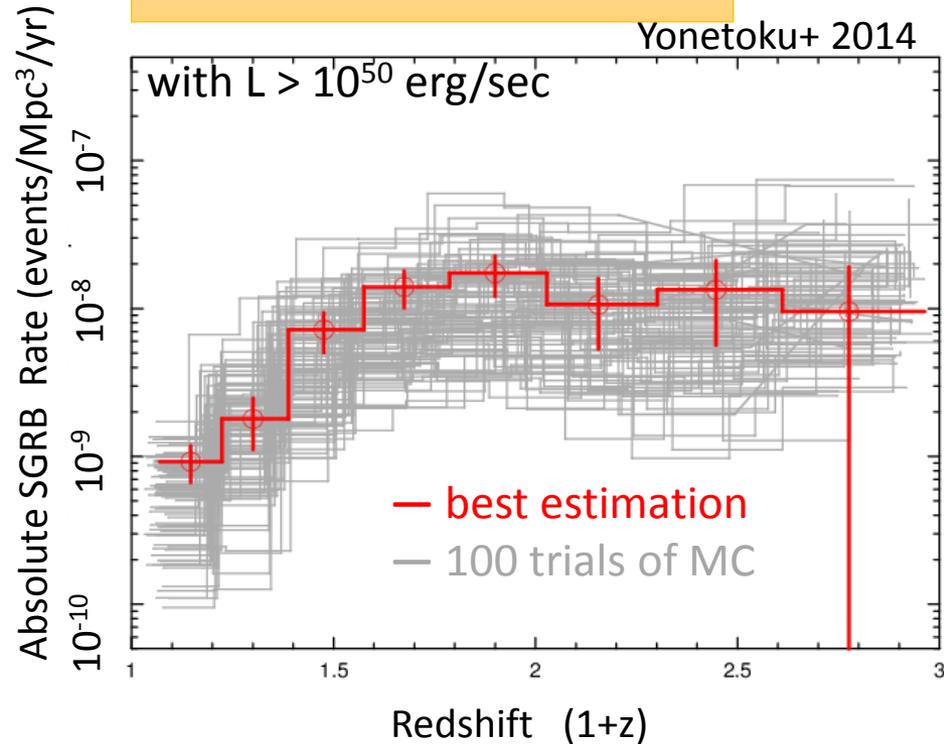
**soft X-ray imager with wide FoV for a microsatellite platform**

# Event rate of SGRBs

## Redshift Distribution



## SGRB Formation Rate



Local Rate including geometrical factor

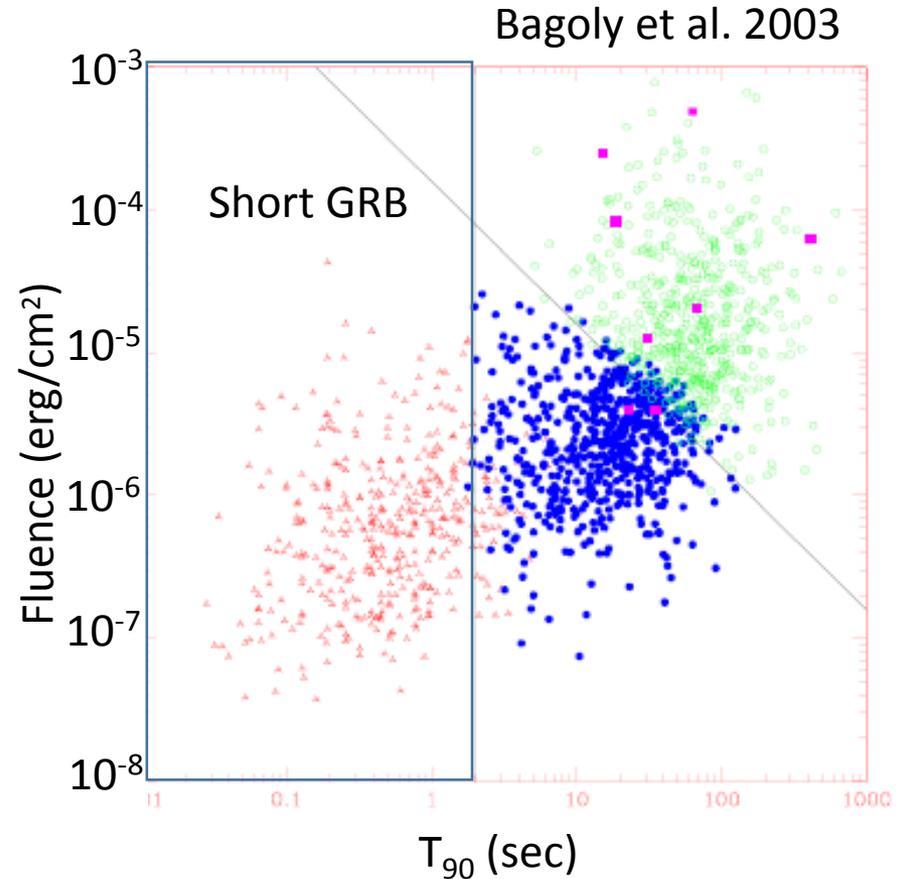
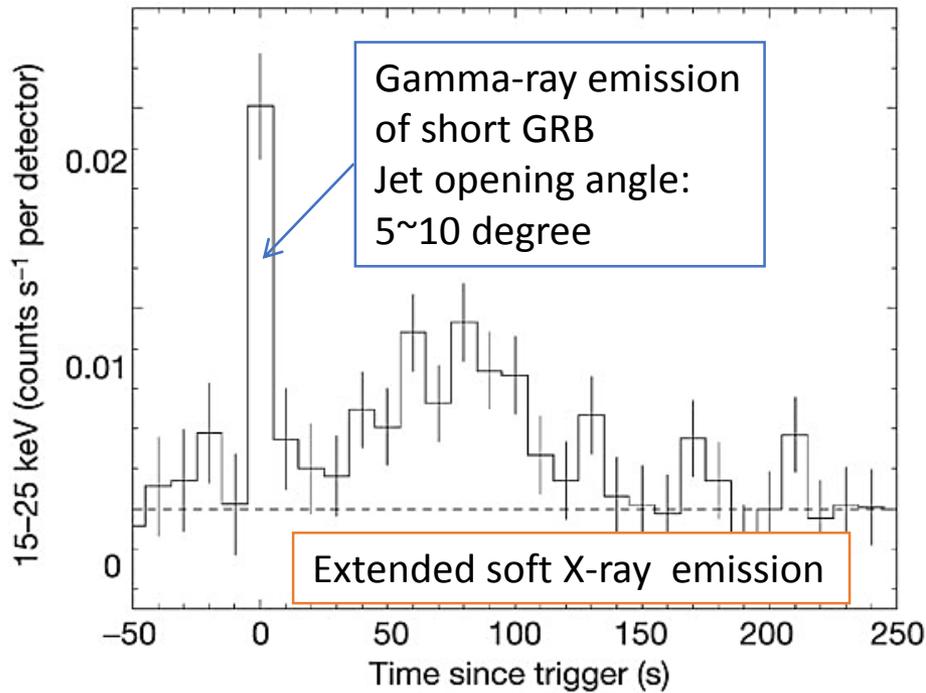
$> 1.15 \times 10^{-7} \text{ events/Mpc}^3/\text{yr}$  (Lower Limit)

$> 3.8 \text{ event/year}$  in  $(200\text{Mpc})^3$  (Lower Limit)

expected to detect GW

$\sim 10 \text{ events/year}$  in  $(200 \text{ Mpc})^3$

# Requirements for Detector



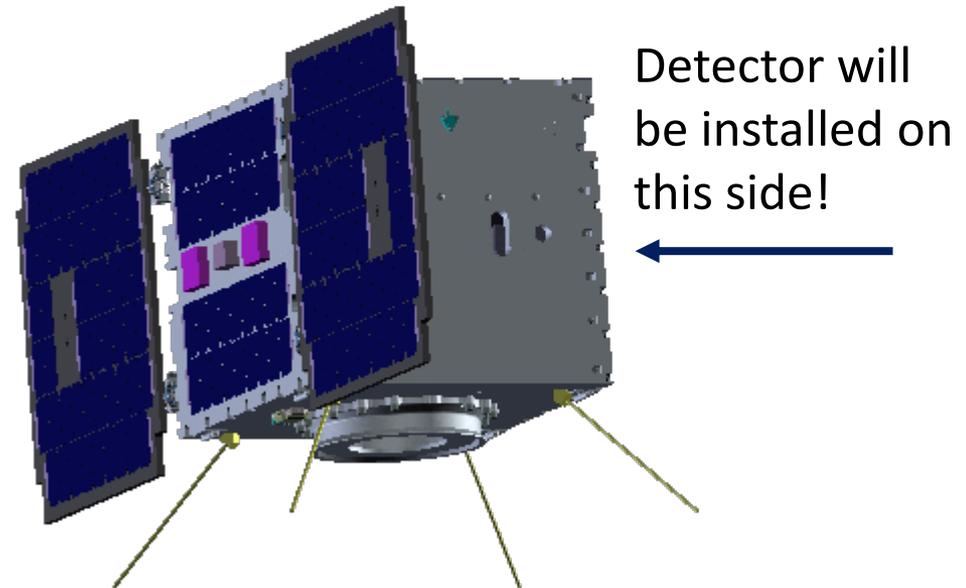
Extended Emission ( $\propto E^{-2}$ , 1~10 keV):

Expected photon flux :  $10^{-6}$  erg/cm<sup>2</sup>  $\sim$  300 photon/cm<sup>2</sup>

Extended emission of nearby SGRBs is enough bright.  
We can observe them by small instruments with 100 cm<sup>2</sup>.

# Micro-satellite proposed from Kanazawa Univ. (Kanazawa-SAT<sup>3</sup>)

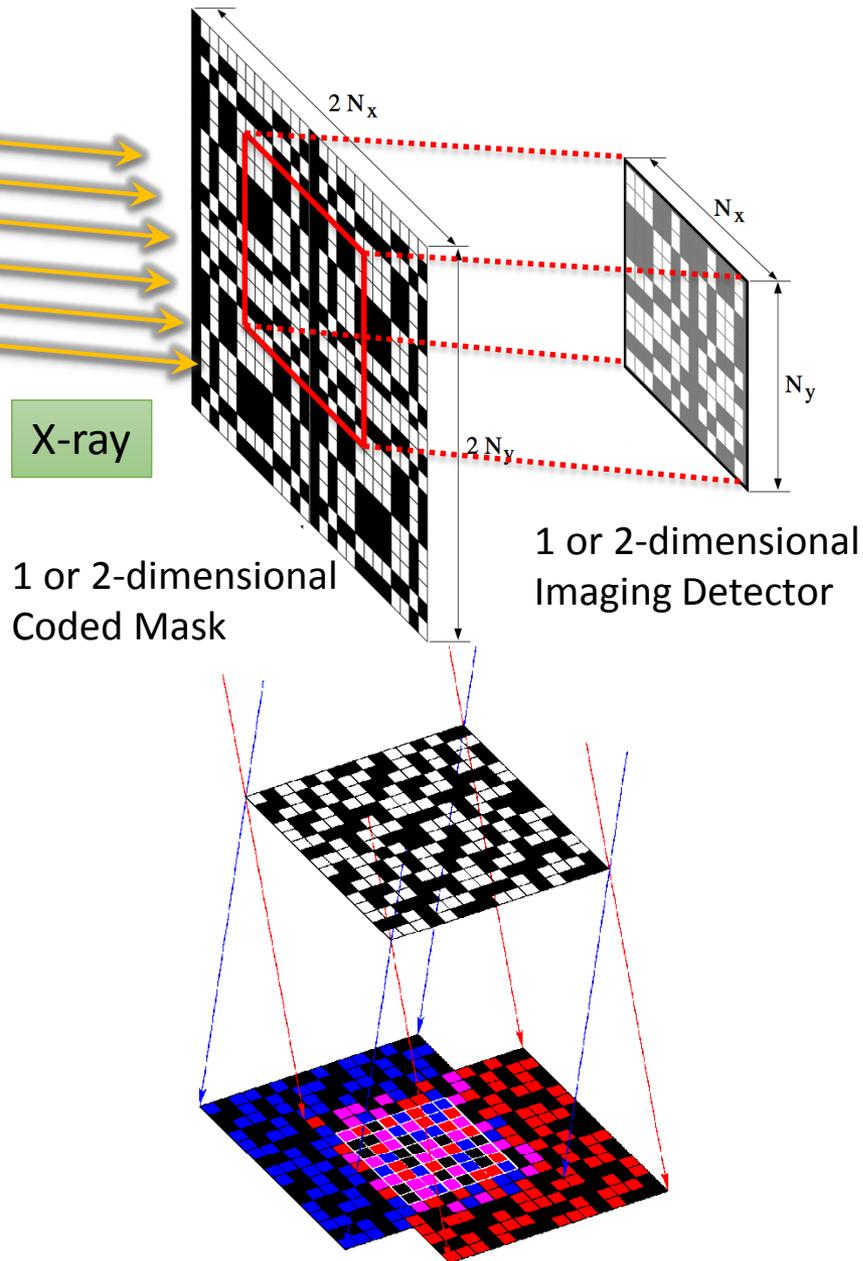
	Kanazawa-SAT <sup>3</sup>
Orbit	LEO
Size	50 × 50 × 50 cm <sup>3</sup>
Weight	< 50 kg
Electric Power	8 W for payload
Telemetry & Command	S-band (downlink) UHF (uplink)



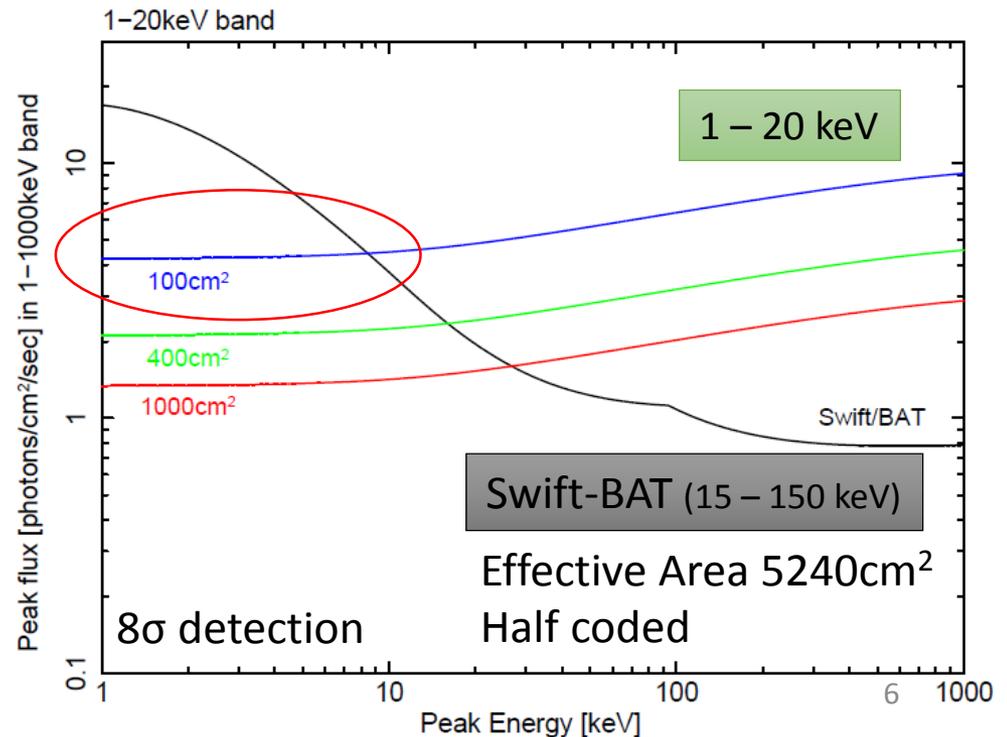
We plan to launch Kanazawa-SAT<sup>3</sup> in 2018-2019!

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# X-ray Imaging Detector



X-ray Imaging Detector	
Detector	Si single side strip ( X and Y)
Energy	1 - 20 keV
Size	0.3mm pitch coded mask 6.4 cm × 7.68 cm × 2
Effective Area	100 cm <sup>2</sup> @ ~1 keV (Half Coded)
Spatial Resolution	~15 arcmin (geometrical)
Field of View	~ 1 str
Weight	5 - 10 kg
Power Consumption	~ 8 W



# Silicon Strip Detector(SSD)

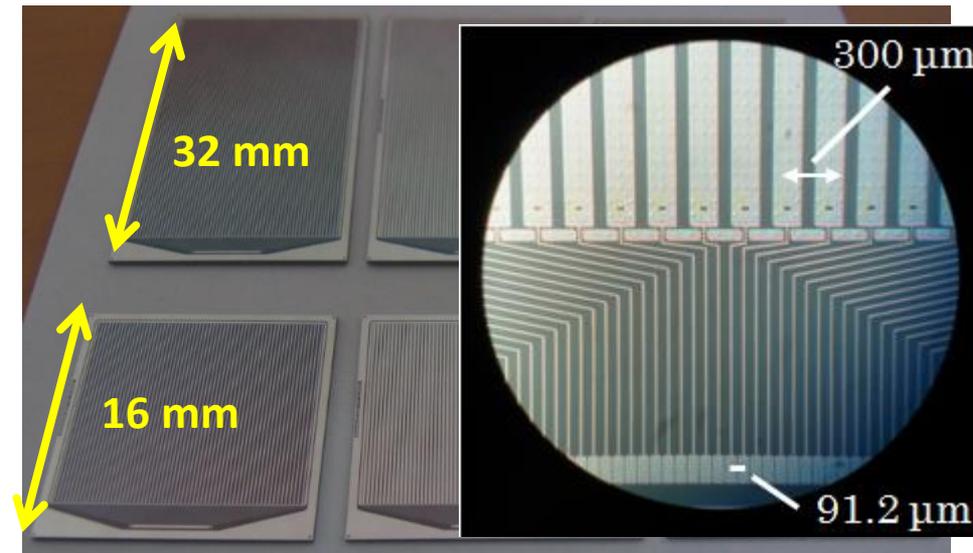
Si strip sensor (Hamamatsu Photonics)

- 64 strips (**0.3mm pitch**)
- 0.5mm thickness (1~20 keV)
- Fan-out structure to connect ASIC
- Capacitance 5 ~ 10 pF each strip

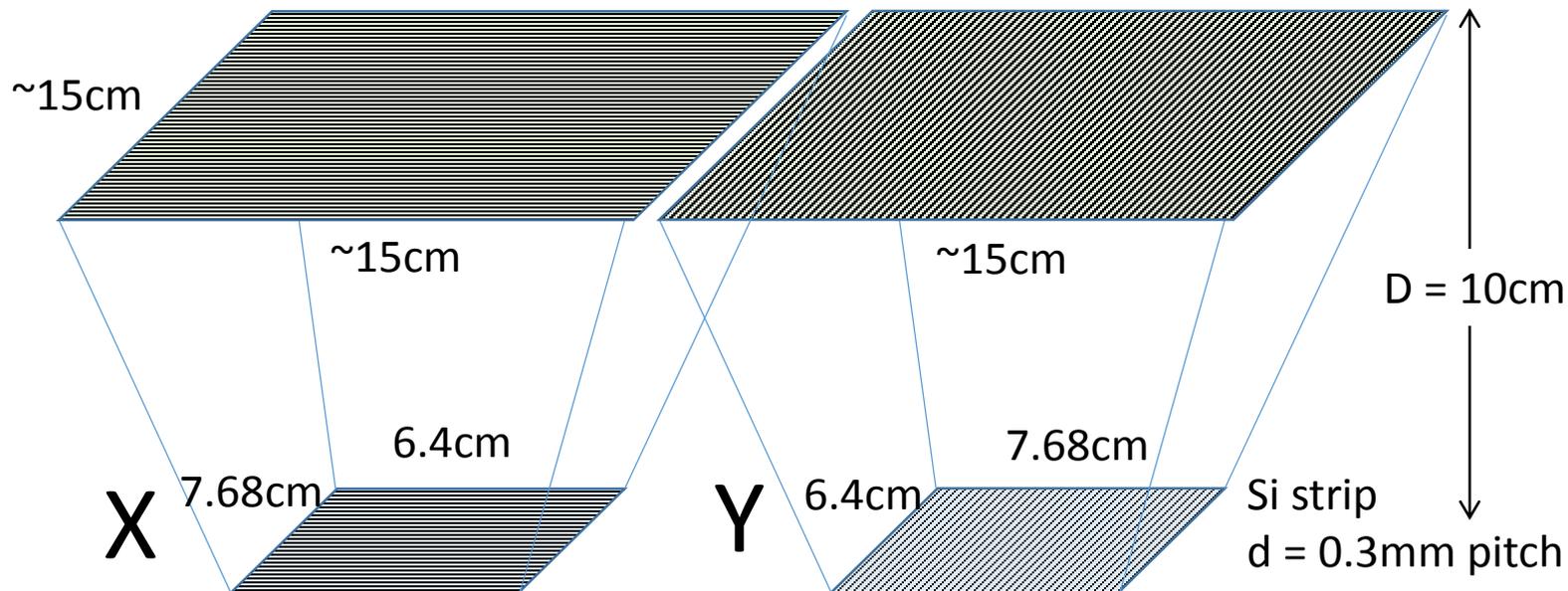
Angular resolution

$$\theta = \tan^{-1}(d/D) = \mathbf{10 \text{ arcmin}}$$

for  $d = 0.3\text{mm}$ ,  $D = 10\text{cm}$



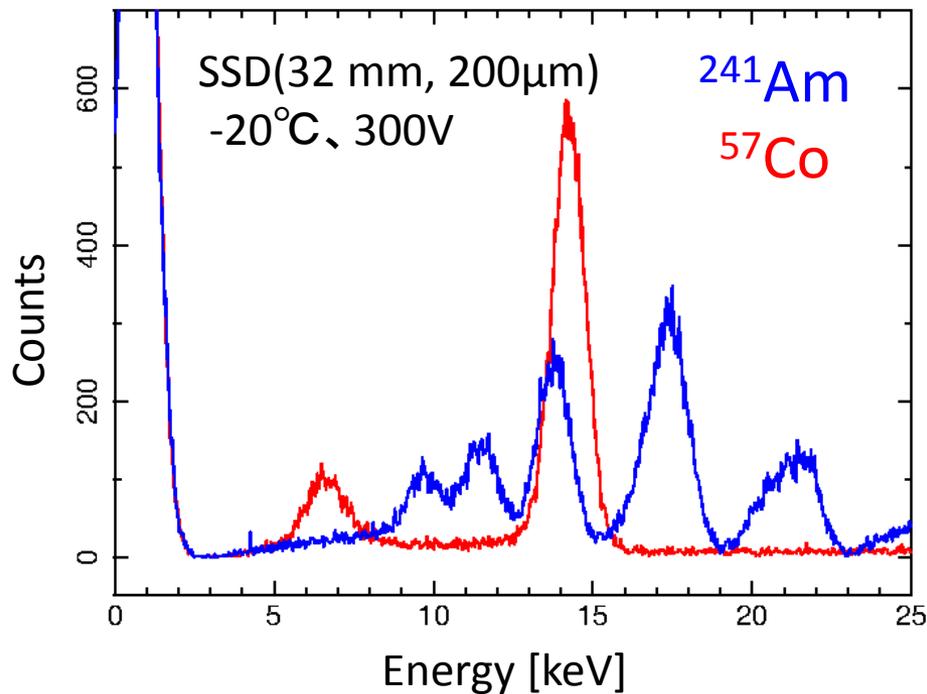
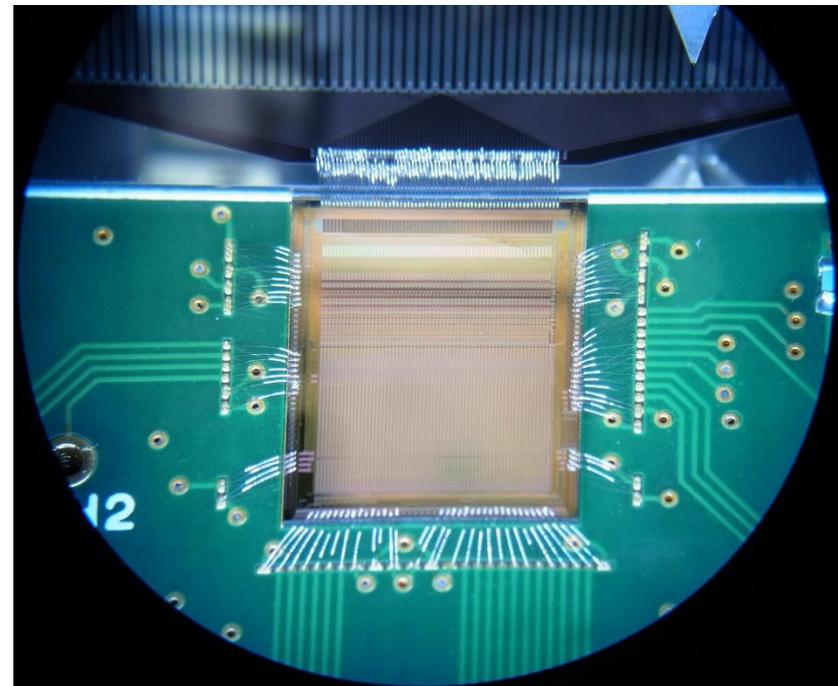
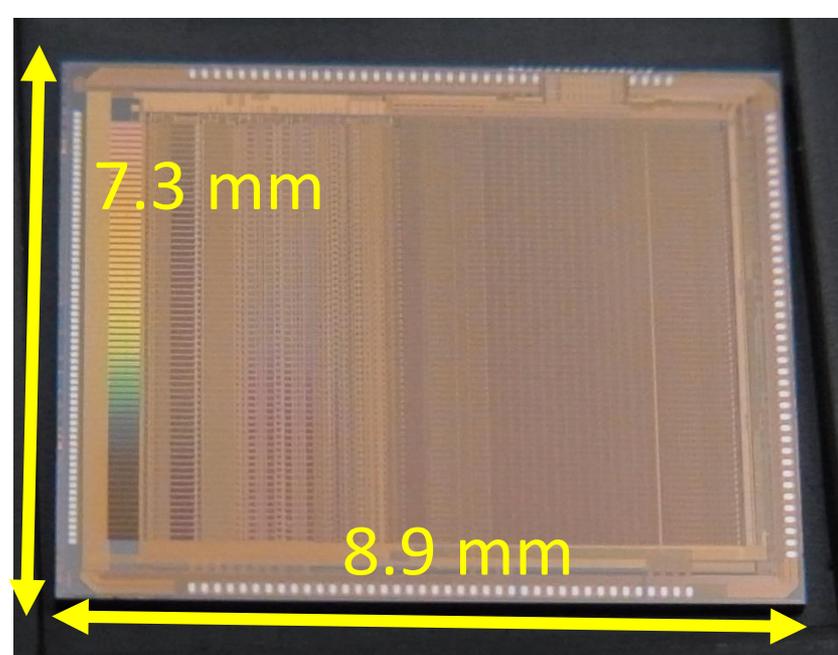
1-dimensional random mask



# Front-end ASIC “ALEX”

- 64-channel analog inputs
- very high gain of  $\sim 750$  mV/fC
- 10-bit Wilkinson-type ADCs
- Power consumption of  $\sim 120$  mW

with the cooperation of  
Prof. Ikeda, Harayama-san (ISAS/JAXA)  
and Arimoto-san (Tokyo Tech.)



# Imaging Exam. - Setup

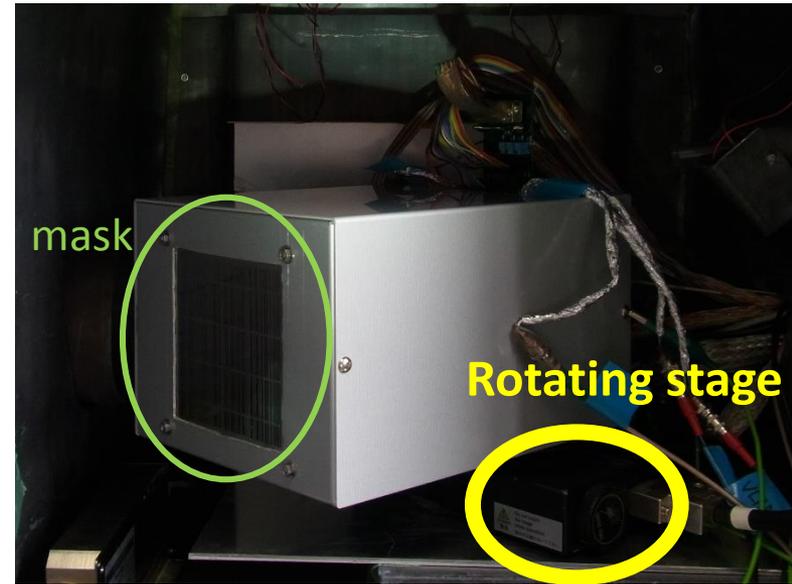
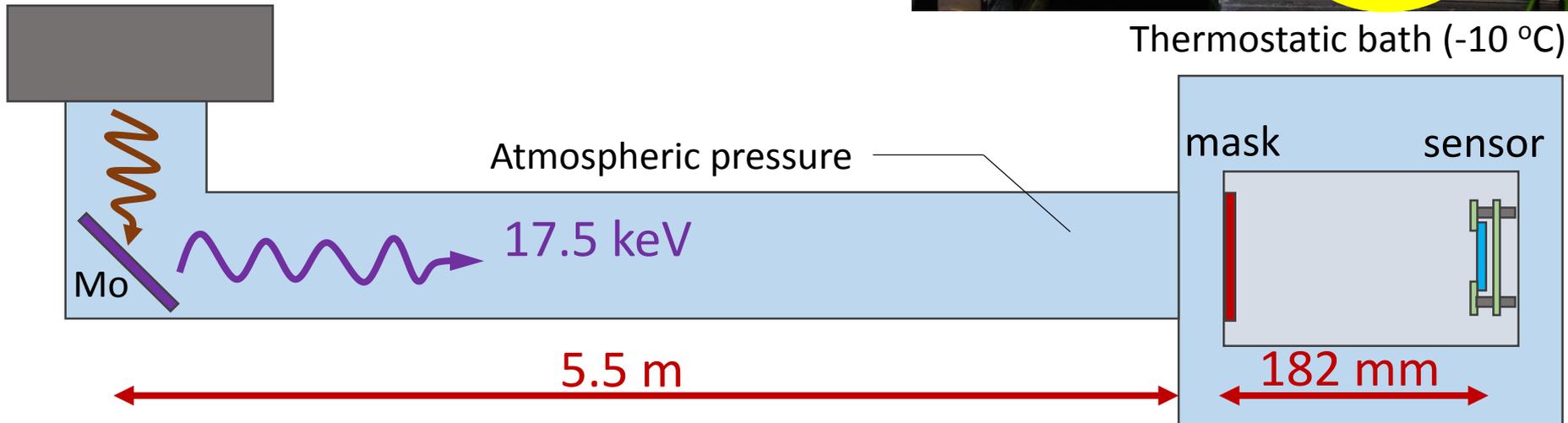
Beamline at lab. of Kanazawa Univ.

- Mono-energy of 17.5 keV produced from molybdenum
- Source divergence  $\sim 20$  arcmin.

## Detector

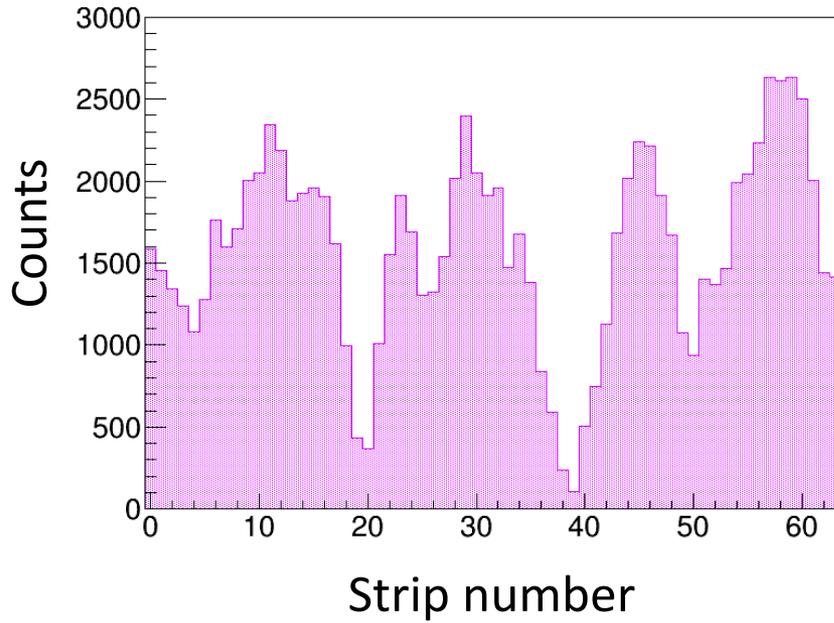
- 64-strip silicon detector
- tungsten mask with a thickness of  $50\ \mu\text{m}$
- Geometrical angular resolution of 5.7 arcmin.

X-ray generator  
(40 kV, 300  $\mu\text{A}$ )



Thermostatic bath ( $-10\ ^\circ\text{C}$ )

# Imaging Exam. - Results



The pattern of the count distribution shifts following the rotating angle.

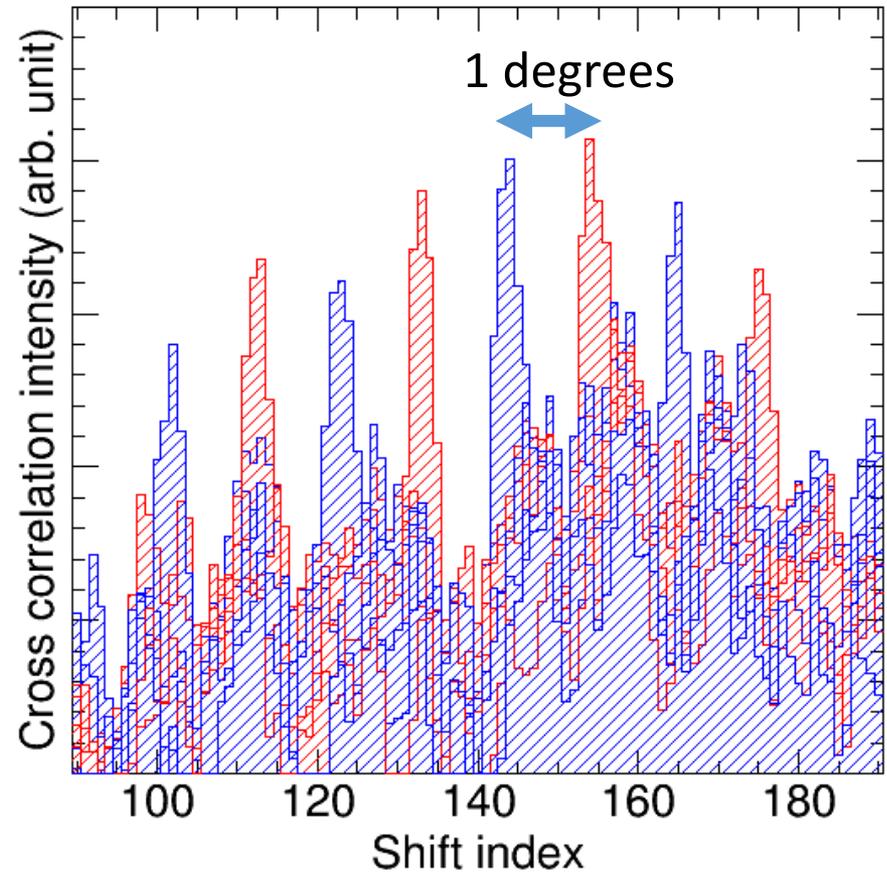
Cross correlation between  
hit and mask patterns

$$C_j = \sum_i d_i \cdot m_{i+j}$$

The equation shows the cross-correlation function  $C_j$  as a sum over  $i$  of the product of the hit pattern  $d_i$  and the mask pattern  $m_{i+j}$ . Red circles highlight  $C_j$ ,  $d_i$ , and  $m_{i+j}$ . Red arrows point from the labels below to these terms.

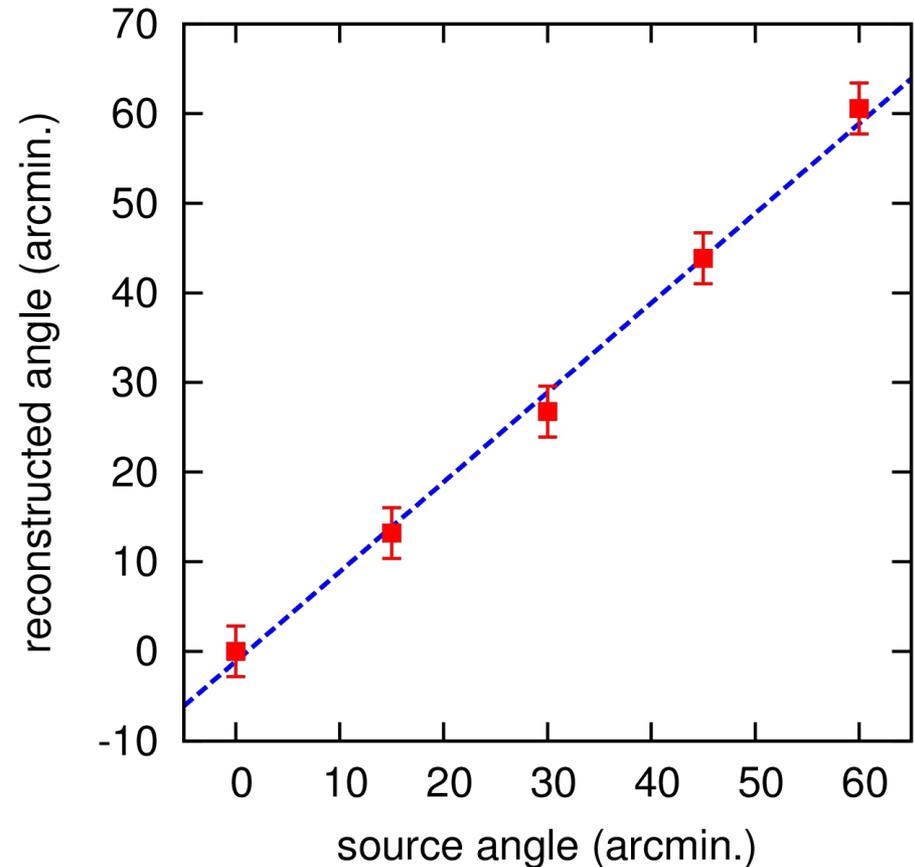
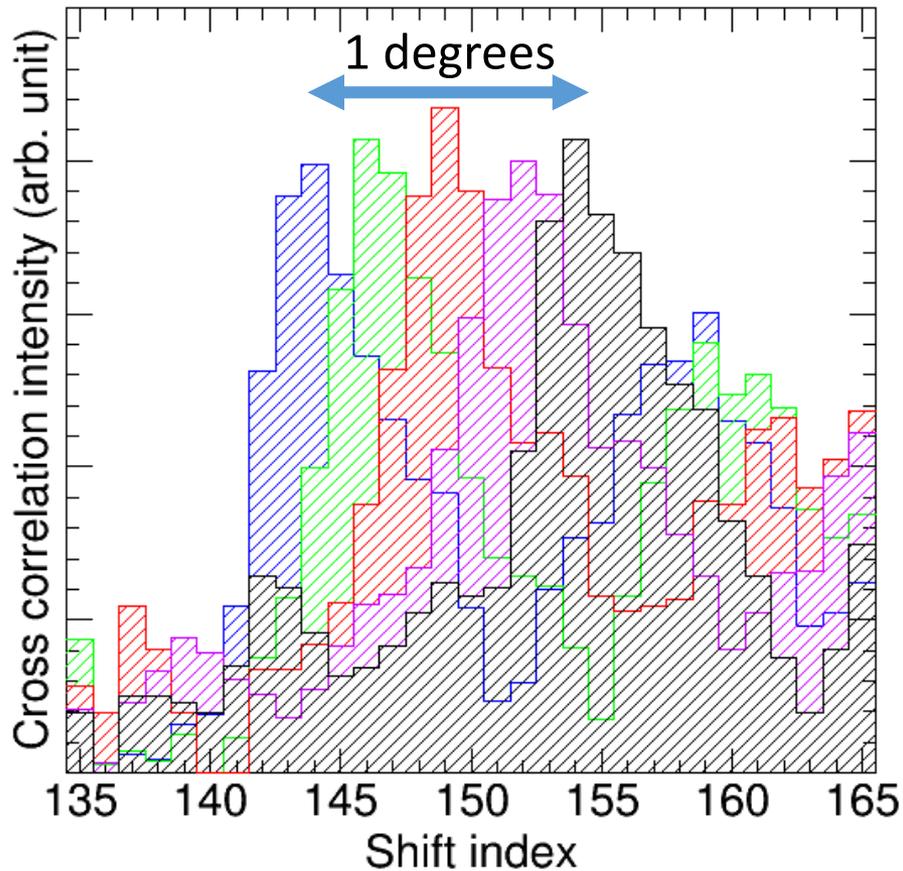
Shift index      hit pattern      mask pattern

8 incident angles  
with an interval of 1 degrees.



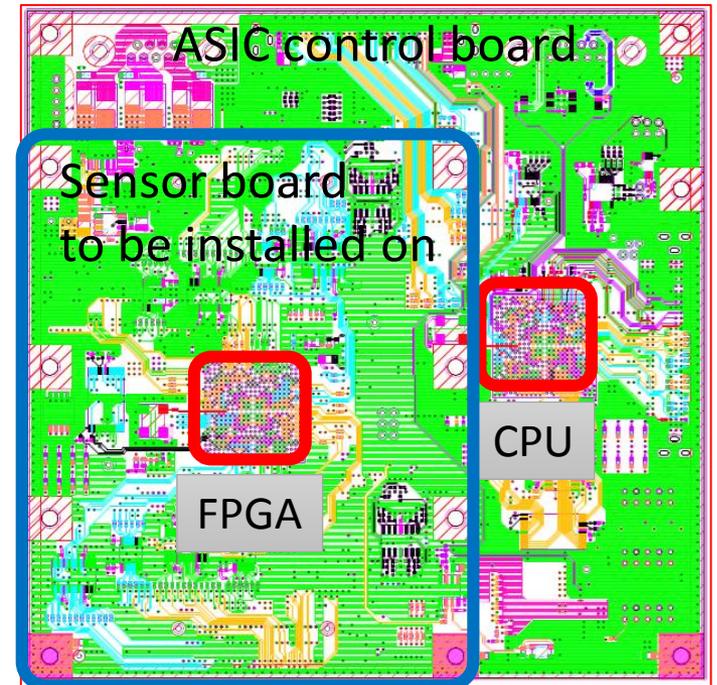
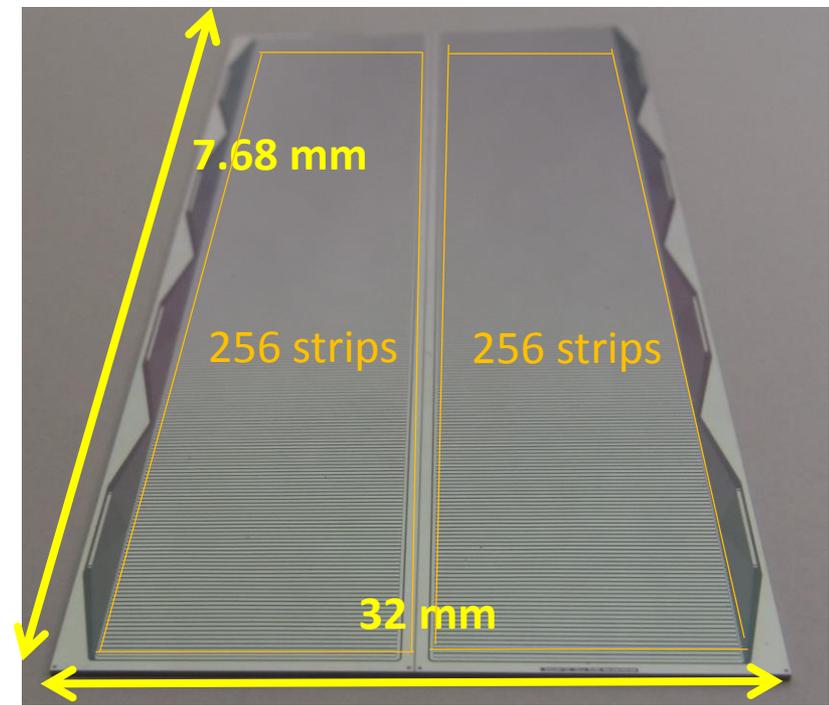
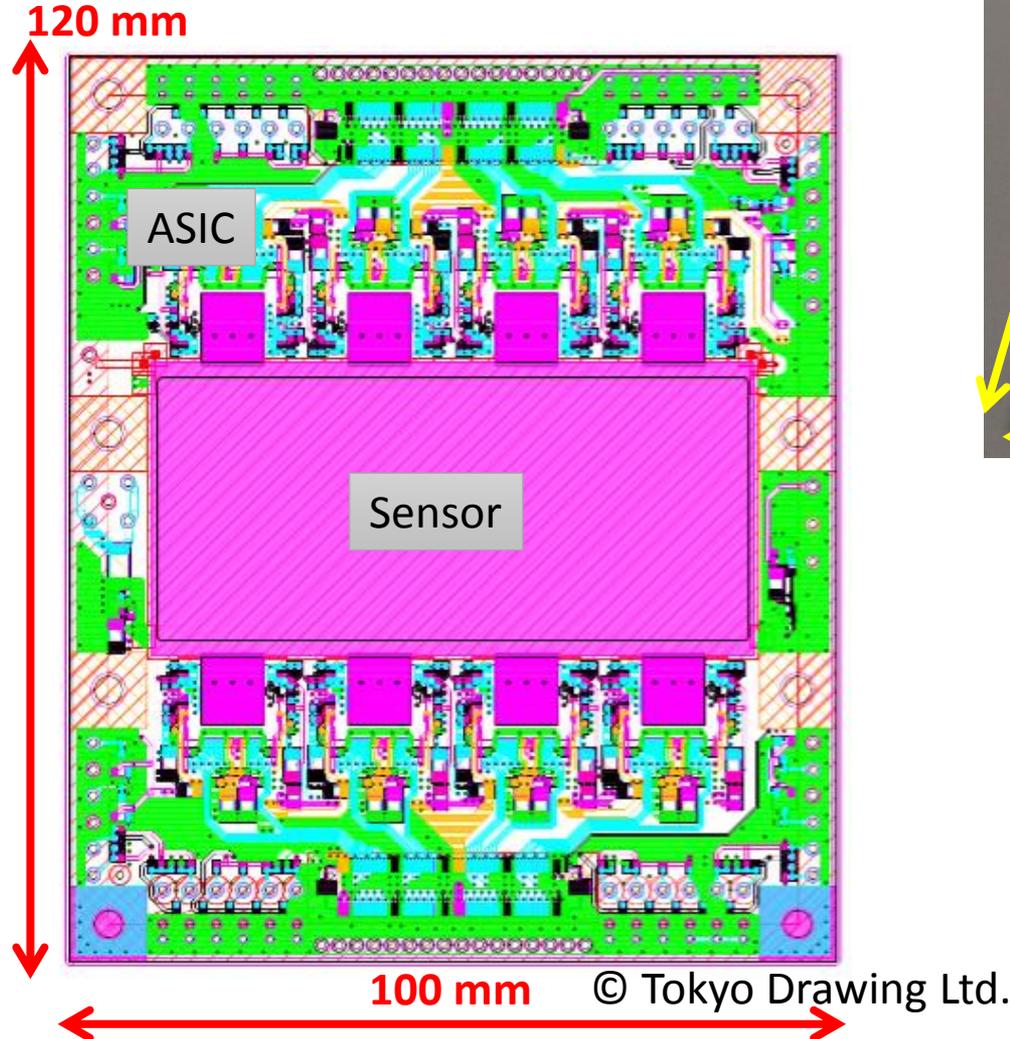
# Imaging Exam. - Results

5 incident angles  
with an interval of 15 arcminutes.



Confirmed the position determination accuracy is consistent with the one determined by the geometric factor of 5.7 arcminutes.

# Prototype Model



Sensor with (256 + 256) strips and 8 ASICs to be installed. ASICs are controlled in parallel.

# Summary

- To detect short GRBs associated with gravitational waves, we are developing a coded X-ray imaging detector consisting with a silicon strip sensor and a tungsten mask.
- We have developed an ASIC for readout of low energy X-rays and confirmed that the detector can measure X-rays from 2 keV to 20 keV.
- We performed the imaging test with a miniature camera system and confirmed the position determination accuracy is consistent with the one determined by the geometrical factor of **5.7 arcminutes**.
- The prototype model for the Kanazawa-SAT<sup>3</sup> is in development, which has a half geometrical area of the flight model, and these technology can be applied to the future HiZ-GUNDAM mission.